

Uitwerking tentamen Golven en Optica 20/3/97 (zonder garantiev)

1a) $V_0 \cos \omega t = \frac{q}{C} + RI + L \frac{dI}{dt} = \frac{q}{C} + R \frac{dq}{dt} + L \frac{d^2q}{dt^2} \Rightarrow \frac{d^2q}{dt^2} + \frac{R}{L} \frac{dq}{dt} + \frac{1}{LC} q = \frac{V_0}{L} \cos \omega t$

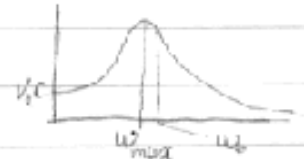
b) $z = A e^{i(\omega t + \alpha)} \Rightarrow (-\omega^2 A + i\omega \frac{R}{L} A + \frac{1}{LC} A) e^{i(\omega t + \alpha)} = \frac{V_0}{L} e^{i\omega t}$ met $\omega_0^2 = \frac{1}{LC}$

$\Rightarrow (\omega_0^2 - \omega^2) A = \frac{V_0}{L} \cos \omega t$ en $\omega \frac{R}{L} A = \frac{V_0}{L} \sin \omega t$

mbv $\sin^2 \omega t + \cos^2 \omega t = 1 \Rightarrow A(\omega) = \frac{V_0/L}{[(\omega_0^2 - \omega^2)^2 + (\frac{R}{L}\omega)^2]^{\frac{1}{2}}}$

$\frac{dA}{d\omega} = -\frac{1}{2} \frac{1}{L \dots} (2(\omega_0^2 - \omega^2) \cdot 2\omega + \frac{R^2}{L^2} \omega) = 0$

$\Rightarrow -2\omega_0^2 + 2\omega^2 + \frac{R^2}{L^2} = 0 \Rightarrow \omega_{\text{max}} = \sqrt{\omega_0^2 - \frac{R^2}{2L^2}} \Rightarrow A_{\text{max}} = \frac{V_0/L}{[\omega_0^2 - R^2/4L^2]^{\frac{1}{2}}}$



c) $R=0 \Rightarrow A(\omega) = \frac{V_0/L}{\omega_0^2 - \omega^2} \Rightarrow$ resonantie als $\omega = \omega_0 = \sqrt{\frac{1}{LC}}$

d) $\gamma = \frac{R}{L} \quad Q = \frac{\omega_0}{\gamma} = \frac{\omega_0 L}{R} = \frac{1}{R} \sqrt{\frac{L}{C}} \quad \omega_{\text{max}} = \omega_0 \sqrt{1 - \frac{R^2}{2L^2 \omega_0^2}} = \omega_0 \sqrt{1 - \frac{1}{2Q^2}}$
 $A_{\text{max}} = \frac{V_0/(L\omega_0 R)}{\sqrt{1 - \frac{R^2}{2L^2 \omega_0^2}}} = V_0 C Q \frac{1}{\sqrt{1 - \frac{1}{2Q^2}}}$ want $\frac{1}{\omega_0 R} = \frac{\sqrt{LC}}{R} = C \frac{L}{\sqrt{LC} R} = C \frac{1}{\sqrt{LC} R} = C \frac{1}{\omega_0 R} = C Q$

Q is omgekeerd evenredig met relatieve breedte resonantiecurve ($Q \approx \frac{\omega_0}{2\Delta\omega}$); ook: voor ongedempt systeem gaat amplitude naar ∞ voor ca. $\frac{Q}{\pi}$ perioden. Grote Q \rightarrow geringe demping, scherp resonantie

e) R dissipatie, C ~ potentiële energie als in veer, L zorgt voor traagheid, vgl kinetische energie massa. Energie opgeslagen in mag.veld L \leftrightarrow electr.veld C.

2. a) $m \frac{d^2x_1}{dt^2} = -kx_1 + k(x_2 - x_1) \quad m \frac{d^2x_2}{dt^2} = -k(x_2 - x_1)$

b) $x_1 = X_1 \cos \omega t, \quad x_2 = X_2 \cos \omega t \Rightarrow -m\omega^2 X_1 = -kX_1 + k(X_2 - X_1) \Rightarrow \frac{X_2}{X_1} = 2 - \frac{m}{k} \omega^2$ ①

$m\omega^2 X_2 = -k(X_2 - X_1) \Rightarrow \frac{X_2}{X_1} = \frac{1}{1 - \frac{m}{k} \omega^2}$ ② $\Rightarrow \frac{m}{k} \omega^2 = 2 - \frac{m}{k} \omega^2 \Rightarrow \omega^2 = \frac{k}{2m} (3 \pm \sqrt{5})$

c) $\frac{X_2}{X_1} = 2 - \frac{m}{k} \omega^2 \Rightarrow \omega_1^2 = \frac{k}{2m} (3 - \sqrt{5}) \Rightarrow \frac{X_2}{X_1} = \frac{1}{2} + \frac{1}{2}\sqrt{5} \quad (\omega_1^2 < \omega_2^2)$

$\omega_2^2 = \frac{k}{2m} (3 + \sqrt{5}) \Rightarrow \frac{X_2}{X_1} = \frac{1}{2} - \frac{1}{2}\sqrt{5}$

d) Potentiële energie benutte veer $\frac{1}{2} k x_1^2$, onderste $\frac{1}{2} k (x_2 - x_1)^2$

\Rightarrow verhouding energie $P = \frac{(x_2 - x_1)^2}{x_1^2} = \left(\frac{x_2}{x_1} - 1\right)^2$

$\omega_1^2 \rightarrow P_1 = \left(-\frac{1}{2} + \frac{1}{2}\sqrt{5}\right)^2 \approx 0.382, \quad \omega_2^2: P_2 = \left(-\frac{1}{2} - \frac{1}{2}\sqrt{5}\right)^2 \approx 2.618$

3 a) $f_{\text{min}} = 40 \cdot 10^6 + 400 \text{ Hz} \quad f'_{\text{max}} = 40 \cdot 10^6 + 2040 \text{ Hz}; \quad \varphi = 0 \Rightarrow$ geen doppler shift \rightarrow

werkelijke $f = \frac{f_{\text{min}} + f_{\text{max}}}{2} = 40 \cdot 10^6 + 1220 \text{ Hz}$

Alternatief: $f' = f_0 / (1 - \frac{v_0 \sin \varphi}{c})$ wel in f'_{max} voor $\varphi = \frac{\pi}{2}$, f'_{min} voor $\varphi = -\frac{\pi}{2}$, etc.

b) $f'_{\text{max}} = 40 \cdot 10^6 + 2040 = (40 \cdot 10^6 + 1220) / (1 - \frac{v_0}{c}) \Rightarrow \frac{v_0}{c} = \frac{(40 \cdot 10^6 + 2040) - (40 \cdot 10^6 + 1220)}{40 \cdot 10^6 + 2040}$

$v_0 = \frac{820 \cdot 3 \cdot 10^8}{40 \cdot 10^6 + 2040} = 6.15 \text{ km/s}$

4 a) $\gamma(\tau) = \frac{\langle E(t) E^*(t+\tau) \rangle}{\langle |E|^2 \rangle} = \frac{\langle E(t) E^*(t+\tau) \rangle}{\langle |E|^2 \rangle} = \langle e^{i\omega t} e^{i[\varphi(t) - \varphi(t+\tau)]} \rangle =$

$= e^{i\omega \tau} \lim_{T \rightarrow \infty} \frac{1}{T} \int_0^T e^{i[\varphi(t) - \varphi(t+\tau)]} dt$

b) $\gamma(\tau) = \frac{e^{i\omega \tau}}{T_0} \int_0^{T_0} e^{i[\varphi(t) - \varphi(t+\tau)]} dt = \frac{e^{i\omega \tau}}{T_0} \int_0^{T_0-\tau} dt + \frac{e^{i\omega \tau}}{T_0} \int_{T_0-\tau}^{T_0} e^{i\varphi} dt =$

$$\frac{d(t+c)}{dt} = \frac{t_0 - t}{t_0} + \frac{t}{t_0} e^{i\omega t} = \left(\frac{t_0 - t}{t_0} + \frac{t}{t_0} e^{i\omega t} \right) e^{i\omega t}$$

u. 4b) Integreer over veel intervallen T_0 , met Δ random $\rightarrow \langle e^{i\Delta} \rangle = 0$

$$\Rightarrow \gamma(\tau) = \frac{t_0 - \tau}{t_0} e^{i\omega \tau} \quad |\gamma(\tau)| = \frac{t_0 - \tau}{t_0} = \frac{1 - 0.25}{1} = 0.75 \rightarrow$$

\rightarrow partiële coherentie

4c) $\tau = 1.25 \rightarrow$ volledige incoherentie, geen contrast in het interferentiepatroon.

5a) $U_p = C \iint e^{ikz} dA = C \iint e^{ik(z_0 + \delta z)} dA = C' \iint e^{ik\delta z} dA = C' \iint g(x,y) e^{ik\delta z} dx dy$



$$\frac{\delta z_x}{x} \approx \frac{x}{L} \quad \delta z_x = \frac{x}{L} x$$

$$\delta z = \delta z_x + \delta z_y = \frac{x}{L} x + \frac{y}{L} y$$

$$\Rightarrow k\delta z = \frac{kX}{L} x + \frac{kY}{L} y = \mu x + \nu y \quad \text{met } \mu = \frac{kX}{L} \text{ en } \nu = \frac{kY}{L}$$

b) $U = \int_{-a/2}^{a/2} dx \int_{-b/2}^{b/2} dy e^{i(\mu x + \nu y)} = \left[\frac{1}{i\mu} e^{i\mu x} \right]_{-a/2}^{a/2} \left[\frac{1}{i\nu} e^{i\nu y} \right]_{-b/2}^{b/2} = \frac{\sin(a\frac{\mu}{2})}{\frac{\mu}{2}} \frac{\sin(b\frac{\nu}{2})}{\frac{\nu}{2}}$

$$I = |U|^2 = \frac{\sin^2(a\frac{\mu}{2})}{(\frac{\mu}{2})^2} \frac{\sin^2(b\frac{\nu}{2})}{(\frac{\nu}{2})^2} \quad \text{met } \mu = \frac{k}{L} X = \frac{2\pi}{\lambda L} X \quad \nu = \frac{2\pi}{\lambda L} Y$$

c) $U = U_x U_y \quad U_x = \int_{-a/2}^{a/2} \cos\left(\pi \frac{x}{a}\right) e^{i\mu x} dx = \int_{-a/2}^{a/2} \frac{e^{i(\pi \frac{x}{a} + \mu x)} + e^{i(-\pi \frac{x}{a} + \mu x)}}{2} dx =$

$$= \frac{\sin\left(a\frac{\mu}{2} + \frac{\pi}{2}\right)}{\frac{1}{2}\left(\mu + \frac{\pi}{a}\right)} + \frac{\sin\left(a\frac{\mu}{2} - \frac{\pi}{2}\right)}{\frac{1}{2}\left(\mu - \frac{\pi}{a}\right)} = \cos\left(a\frac{\mu}{2}\right) \left[\frac{1}{\frac{1}{2}\left(\mu + \frac{\pi}{a}\right)} - \frac{1}{\frac{1}{2}\left(\mu - \frac{\pi}{a}\right)} \right]$$

$$U_y = \int_{-b/2}^{b/2} \left(0.5 + 0.5 \cos\left(\frac{2\pi y}{b}\right)\right) e^{i\nu y} dy = \sin\left(b\frac{\nu}{2}\right) \left[\frac{1}{\nu} - \frac{1}{\nu + \frac{2\pi}{b}} - \frac{1}{\nu - \frac{2\pi}{b}} \right]$$

d) Zijlobben kleiner, minder verzwamd bij dubbelsterren waarvan de een veel lichtzwaarder dan de ander.